



# Laser Sounder for Measuring Atmospheric CO<sub>2</sub> Concentrations for the ASCENDS Mission - Progress

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# Abstract

Accurate global measurements of tropospheric CO<sub>2</sub> concentrations with diurnal coverage and monthly temporal resolution are needed to better quantify the processes that exchange atmospheric CO<sub>2</sub> with the land and oceans. The National Research Council's 2007 Decadal Survey for Earth Science recommended following the OCO and GOSAT CO<sub>2</sub> measuring missions with a laser-based mission called ASCENDS.

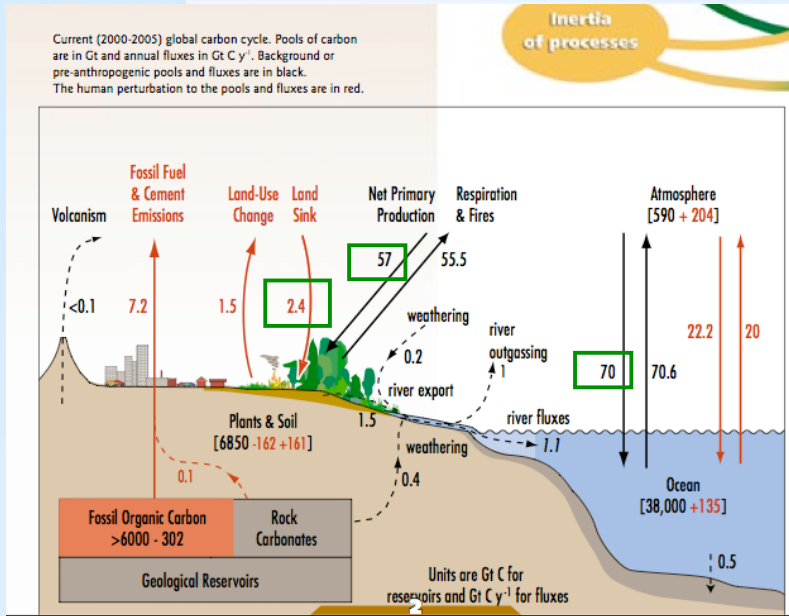
We have been developing and analyzing the laser technique and technologies for the remote measurement of tropospheric CO<sub>2</sub> concentrations from space. Our initial goals are to develop and demonstrate the lidar techniques and technologies that permit measurements of the CO<sub>2</sub> column abundance over horizontal paths and from aircraft at the few-ppmv level. Our longer-term goal is to demonstrate the needed capabilities of the technique and technologies, and develop a space mission approach and the instrument design for the ASCENDS mission. This work is ongoing and has been supported by the NASA ESTO ACT and IIP programs.

Our approach is to use the 1570-nm CO<sub>2</sub> band and a pulsed dual channel laser absorption spectrometer. This uses differential lidar absorption measurement in an altimeter mode, and continuously measures at nadir from a near-polar circular orbit. It uses several tunable fiber laser transmitters allowing simultaneous measurement of the absorption from a CO<sub>2</sub> absorption line in the 1570 nm band, O<sub>2</sub> extinction in the oxygen A-band, as well as surface height and aerosol backscatter in the same measurement path. It directs the narrow co-aligned laser beams toward nadir, and measures the energy of the pulsed laser echoes reflected from land and water surfaces. During the measurement, the lasers are tuned across a selected CO<sub>2</sub> line and a region between two O<sub>2</sub> lines near 765 nm. The lasers have spectral widths much narrower than the gas absorption lines and are wavelength tuned at kHz rates. The receiver uses a telescope and photon counting detectors, and measures the background light and energies of the laser echoes from the surface, along with scattering from any clouds and aerosols in the path. The gas extinction and column densities for the CO<sub>2</sub> and O<sub>2</sub> gases are estimated from the ratios of the on and off line signals. We use pulsed laser signals and time gating to isolate the laser echo signals from the surface, and to reject photons scattered from thin clouds and aerosols in the path, which can otherwise bias retrievals. High signal-to-noise ratios are required and the gas column absorption estimates need to be quite stable for hours.

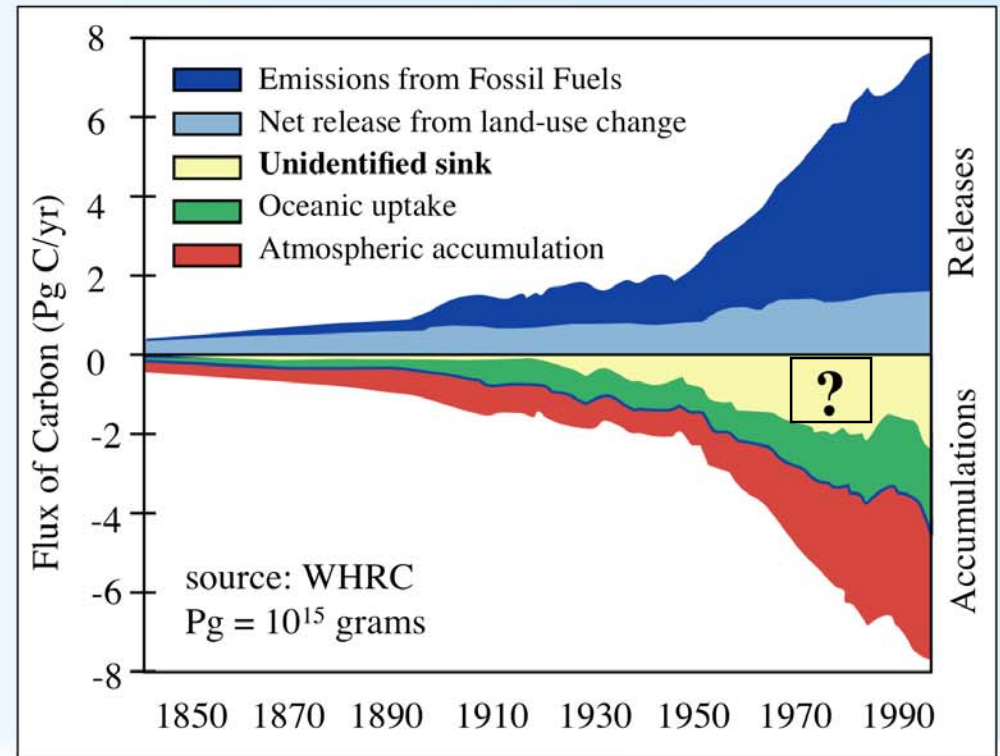
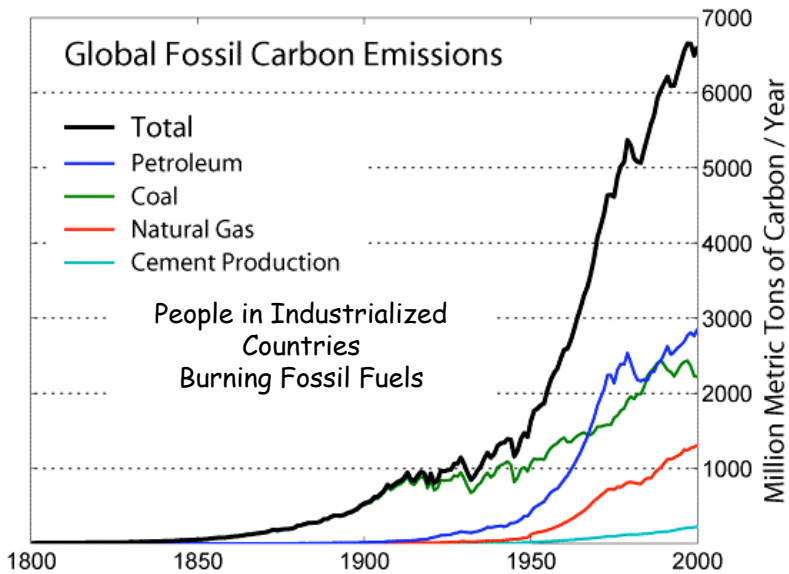
We have constructed breadboard versions of the CO<sub>2</sub> and O<sub>2</sub> sensors, which use fiber lasers and a 20 cm diameter telescopes. We have used them to make measurements of CO<sub>2</sub> and O<sub>2</sub> absorption in the laboratory and over 206, 400-m 1.3 and 2.2 km long open horizontal paths. These have been in several sessions extending over multiple days, and have allowed us to compare its estimates to readings from external CO<sub>2</sub> sensors. We have also calculated several characteristics of the technique for space, including its expected measurement performance, selected some key technologies for space and have performed a space mission accommodation study. We will show these results in the paper.



# Atmospheric CO<sub>2</sub> - Cycle, Sources, Sinks



- Of anthropogenic CO<sub>2</sub> emitted to date, ~ 30% can not be accounted for - the "unknown sink"
- The "unknown sink" may be Northern Hemisphere forests.
- Will this sink continue to operate in the future?
- How will CO<sub>2</sub> fluxes in polar regions respond to warming?





# Laser Sounder for ASCENDS Mission

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## Why lasers ?

- Measures at night & at all times of day
- Continuous "glint" measurements over oceans
- Measures at high latitudes
- Illumination path = observation path
- Smaller measurement footprint
- Measures through broken clouds

## Why CO2 Sounder approach ?

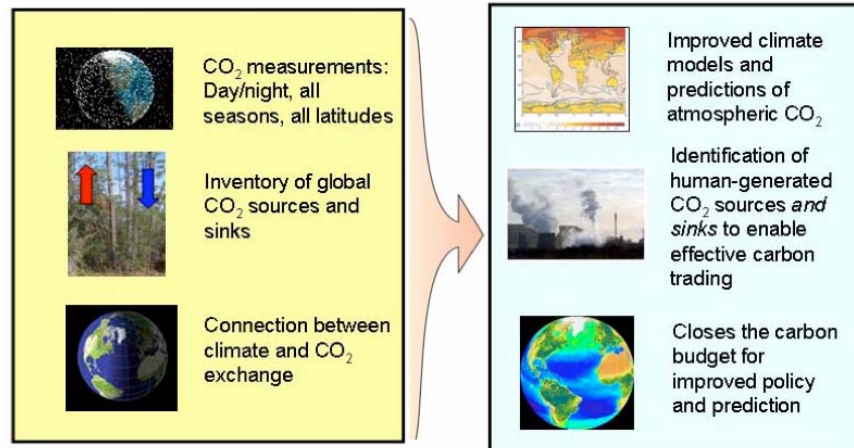
- Excellent CO2 & O2 lines at 1572 & 765 nm
- Weight column ~ uniformly in lower troposphere
- Time gating isolates the full column signal
- Surface height & aerosol backscatter profiles
- Multi-wavelengths minimize instrumental errors
- Strong laser & detector technology bases

## Active Sensing of CO<sub>2</sub> Emissions over Nights, Days, and Seasons (ASCENDS)

Active Sensing of CO<sub>2</sub> Emissions over Nights, Days, and Seasons (ASCENDS)  
 Launch: 2013-2016  
 Mission Size: Medium



*Ascends  
Mission from  
US NRC  
Decadal  
Survey*



Need CO<sub>2</sub> & O<sub>2</sub>  
Lower  
tropospheric  
column  
measurements



# ASCENDS Mission - Laser Sounder Approach

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## 3 simultaneous laser measurements

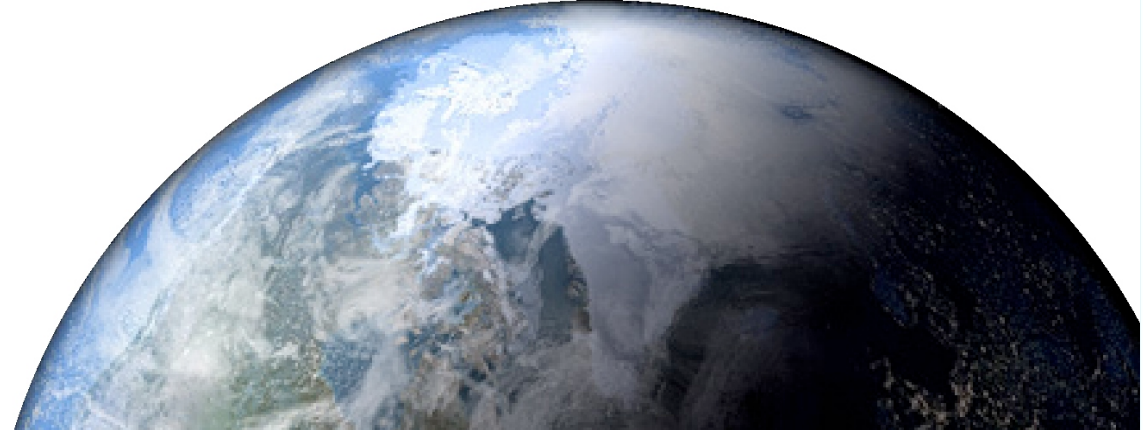
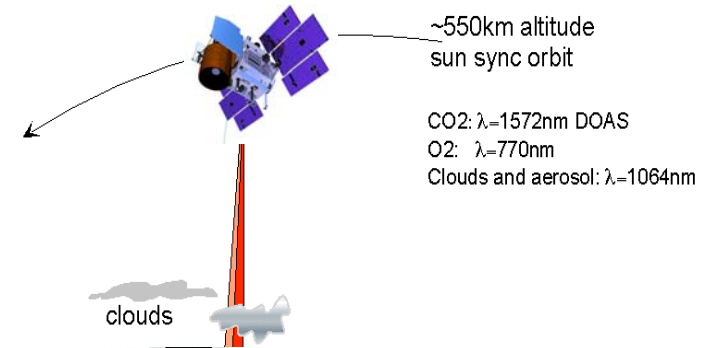
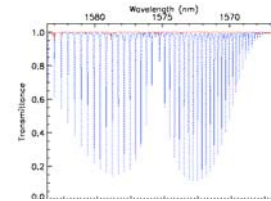
1. CO<sub>2</sub> lower tropospheric column  
One line near 1572 nm
2. O<sub>2</sub> total column  
Measured between 2 lines near 765 nm
3. Altimetry & atmospheric backscatter profile:  
Surface height and atmospheric scattering profile at ~ 1064 nm

### Measurements use:

- Pulsed EDFA lasers
- KHZ pulse rates
- 6 laser wavelengths/ gas line
- Time gated Photon counting receiver

### Measures:

- CO<sub>2</sub> tropospheric column
- O<sub>2</sub> tropospheric column
- Cloud backscattering profile



### CO<sub>2</sub> & O<sub>2</sub> column measurements:

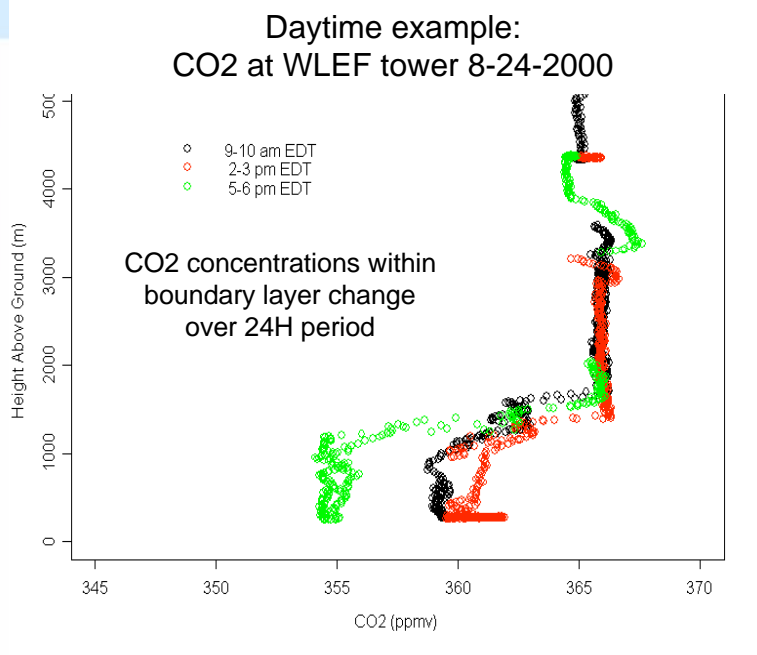
- Pulsed (time gated) signals :
  - Isolates full column signal from surface
  - Reduces noise from detector & solar background
- Target: ~ 1ppmV in ~100 km along track sample



# Dynamic CO<sub>2</sub> concentrations & Orbital Sampling

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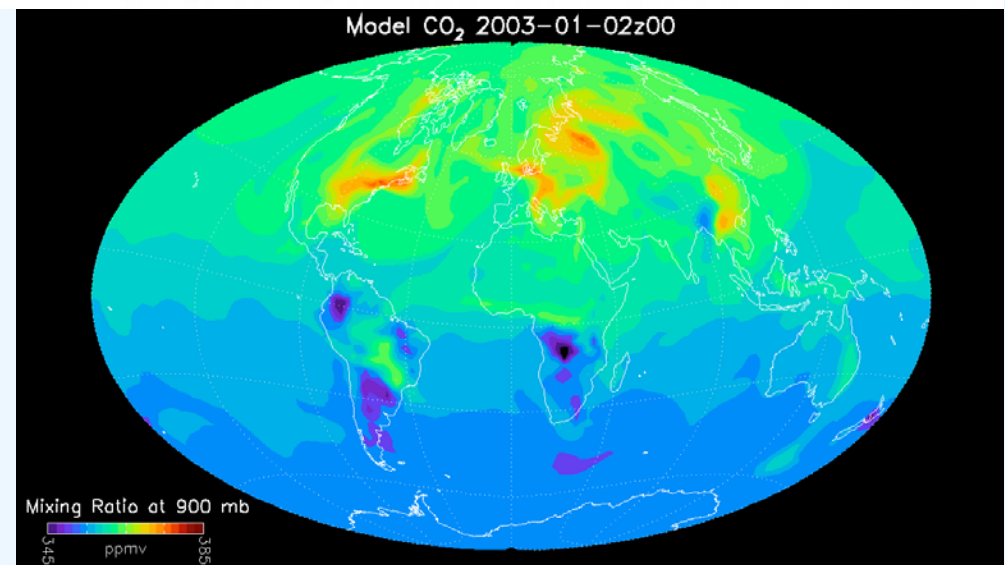
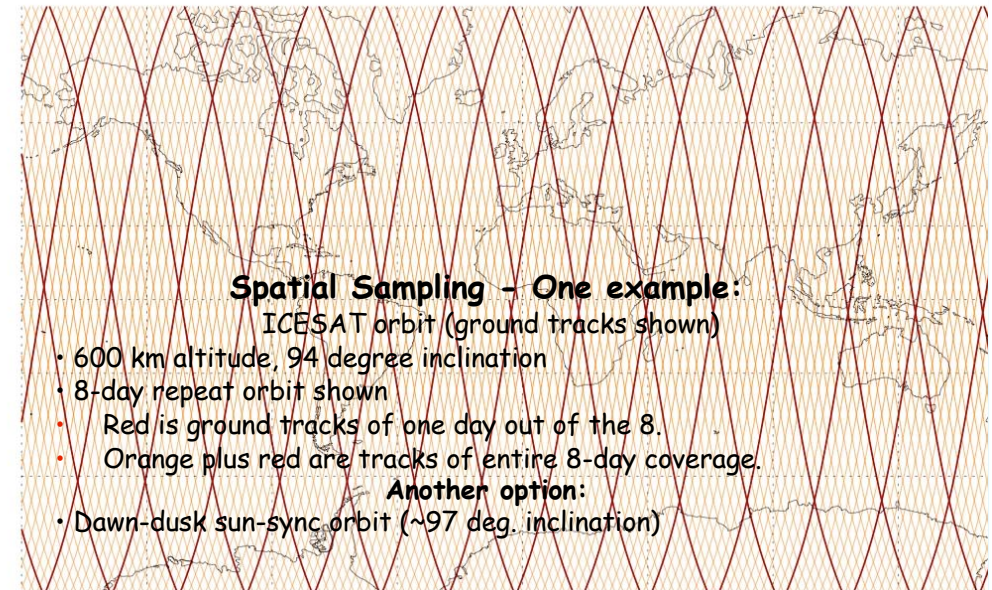
## Vertical (diurnal) variability



## Spatial (x,y) variability

Global CO<sub>2</sub> Concentration Simulation at 900 mbar pressure altitude (Kawa et al., JGR, 2004)

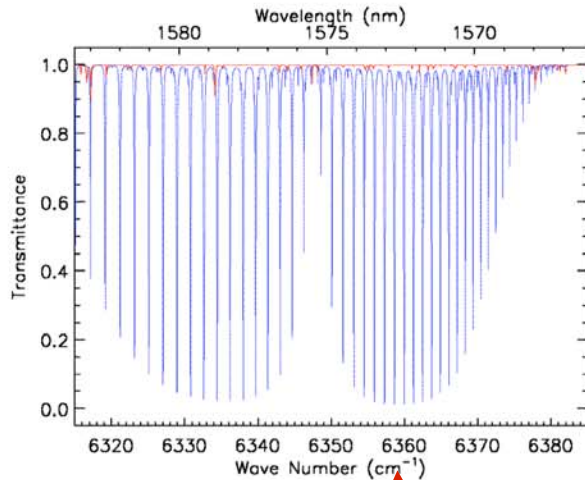
The output step of the movie is 6 hrs, i.e., 4/day.  
The Simulation runs on a 15-min time step.





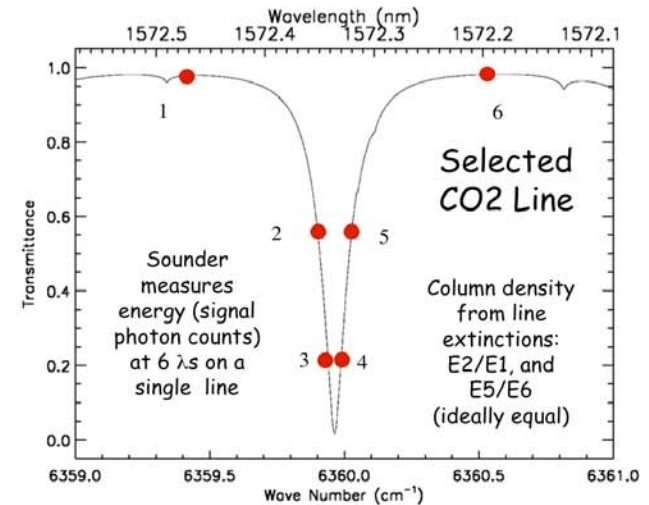
# CO<sub>2</sub> Band & Line Measurement Approach

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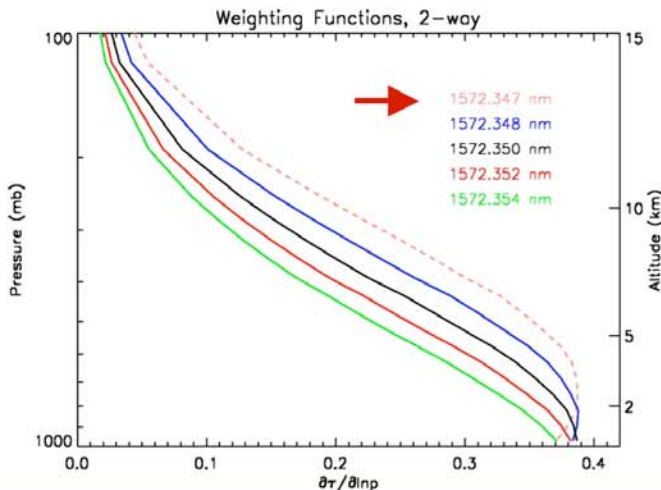
1570 nm CO<sub>2</sub>  
Absorption  
Band from  
Space  
(HITRAN)

Extinction of  
lines vary with  
# of CO<sub>2</sub>  
molecules in  
column



Sequentially step through the  $\lambda$ 's in time

Column Altitude Weighting Function (pts 2 & 5)



Using laser tuned to sides of Absorption line provides Column Measurement weighted to lower trop., via CO<sub>2</sub> Line Broadening

Lasers Provide:

- Narrow measurement line widths (MHz)
- Tunable Stable frequencies (MHz)

Energy Measurement Resolution:

- Need ~ 1000:1 SNR for online energies (E2, E5)
- With similar errors for O<sub>2</sub> gas measurement, results in ~ 1 ppm error in CO<sub>2</sub> mixing ratio



# Gas column retrievals from measurements

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Accurate estimates of N depend on knowledge of :

- $\sigma$  - line cross section
- "z" effective path length
- $\lambda$  - laser wavelengths
- Pref : Transmitted laser power ( $\lambda$ )
- Tsys: System transmission ( $\lambda$ )
- $E_r$  (high SNRs)

Some error sources:

- $\sigma$  - temp effects in line cross section
- z, from atmospheric scattering
- System changes; small  $\lambda$ -dependences in:

$$E_{tr}, \tau_{sys}(\lambda_{on})/\tau_{sys}(\lambda_{off})$$

- Noise (signal & background shot noise, detector noise) in detected echo signal

Goal :

- Maximize received SNR
- Minimize all other error sources

General form of DIAL equation for uniform horizontal path (Beer's Law):

$$\frac{E_r(\lambda_{on})}{E_r(\lambda_{off})} = \frac{E_{tr}(\lambda_{on}) \tau_{sys}(\lambda_{on})}{E_{tr}(\lambda_{off}) \tau_{sys}(\lambda_{off})} \exp(-\sigma N_g z)$$

Estimated CO2 number density:

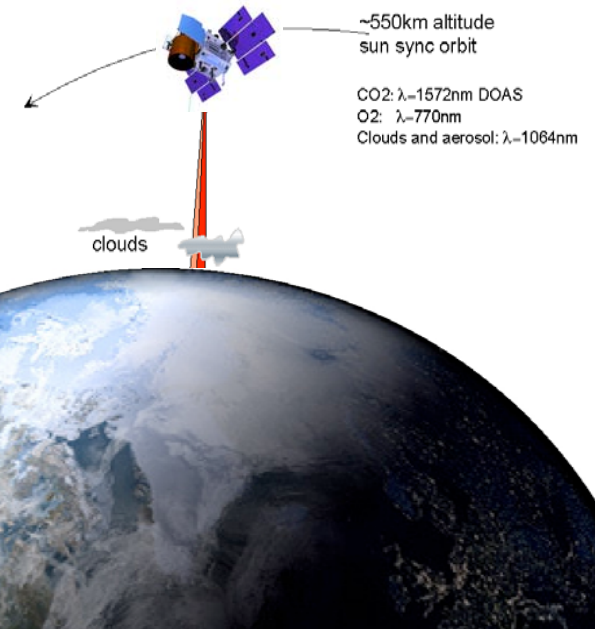
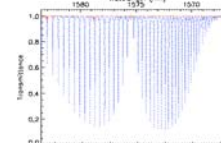
$$N_g = \frac{1}{\sigma z} \ln \left\{ \frac{E_r(\lambda_{off})}{E_r(\lambda_{on})} \frac{E_{tr}(\lambda_{on}) \tau_{sys}(\lambda_{on})}{E_{tr}(\lambda_{off}) \tau_{sys}(\lambda_{off})} \right\}$$

SNR

"Stability"

Measures:

- CO2 tropospheric column
- O2 tropospheric column
- Cloud backscattering profile

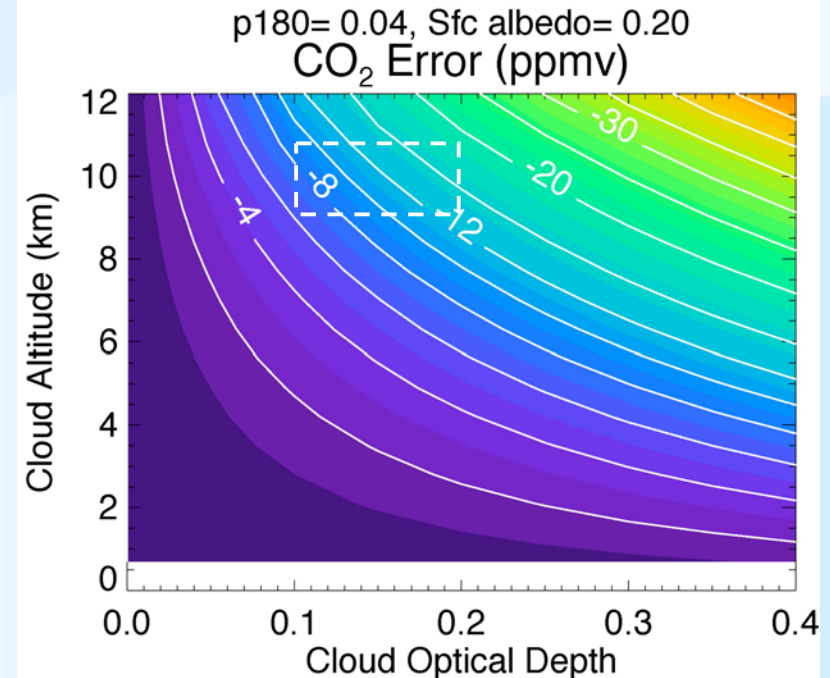
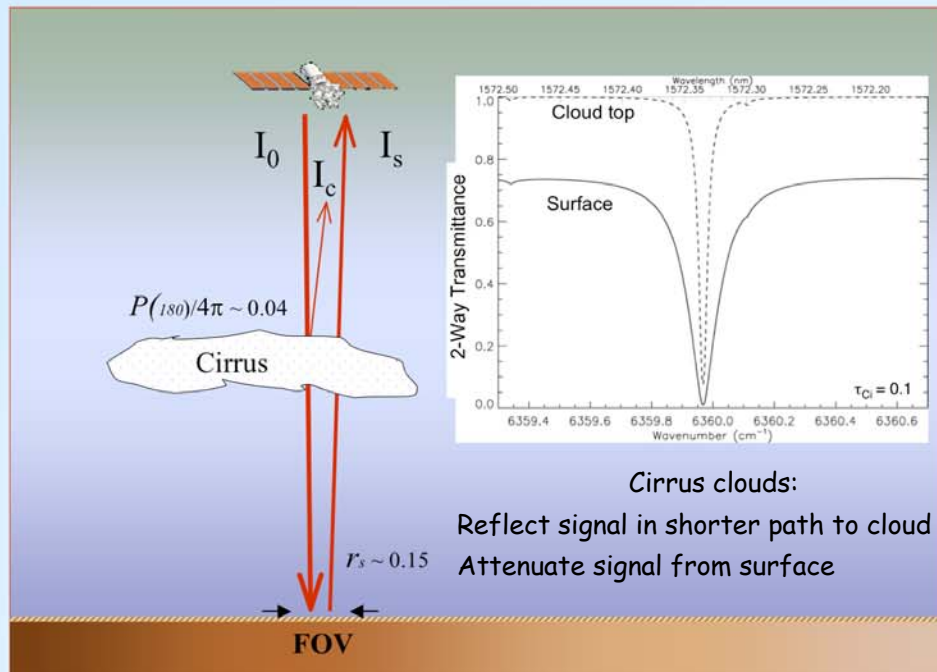






# Atmospheric Scattering - Impact on Non-range gated measurements

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- Cirrus clouds are quite prevalent
- Cloud reflections shorten average optical path → bias CW (non-gated) column estimates
- Cirrus cloud scattering → 8-14 ppm errors in non-range gated measurements
- Errors led our team to use a pulsed (& range gated) approach
- Range gating eliminates cloud scattering errors (except for ground fogs)



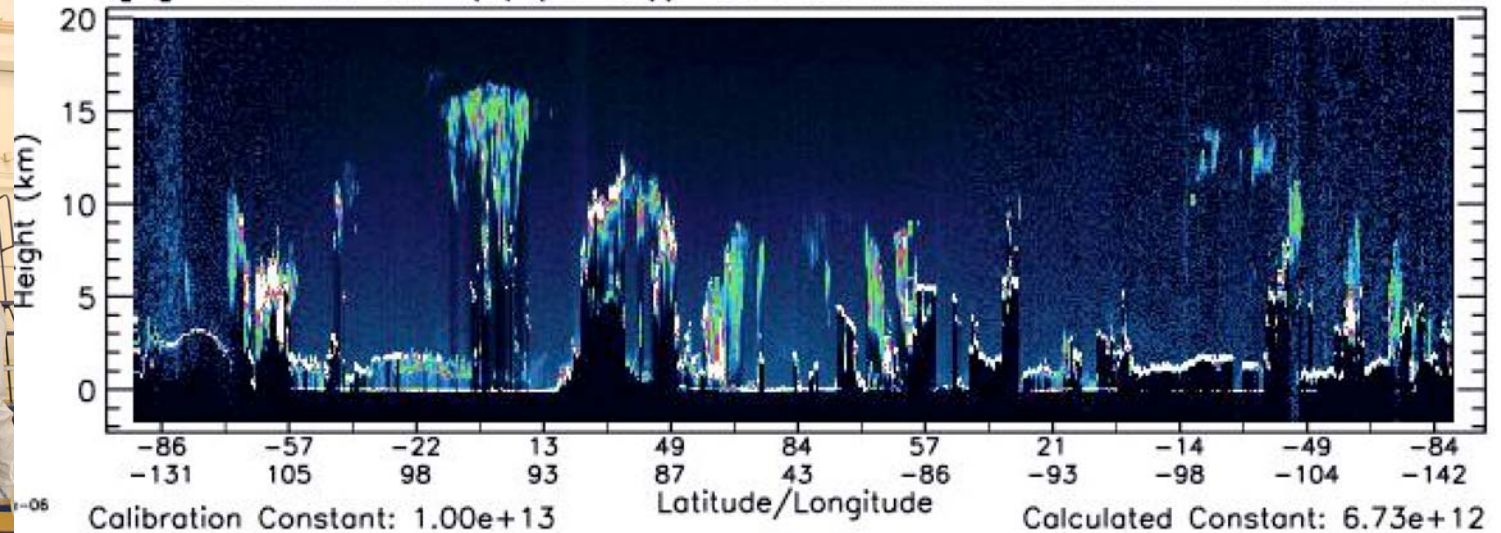
# Atmospheric Scattering & measurement approach

## Atmospheric Backscatter:

- Scattering structure is complex & variable
- Most nadir-zenith paths have some scattering above surface
- Many instances of thin cirrus clouds & aerosols
- Some thick clouds
- "Target Depth" with clouds ~15 km => ~133 usec in travel time

Atmospheric Profile measured by GLAS/ ICESat (from campaign L2a)

[1] 532 nm CAB (1/(m-sr)) 15-Nov-2003 12:00:00 - 13:36:31 GMT





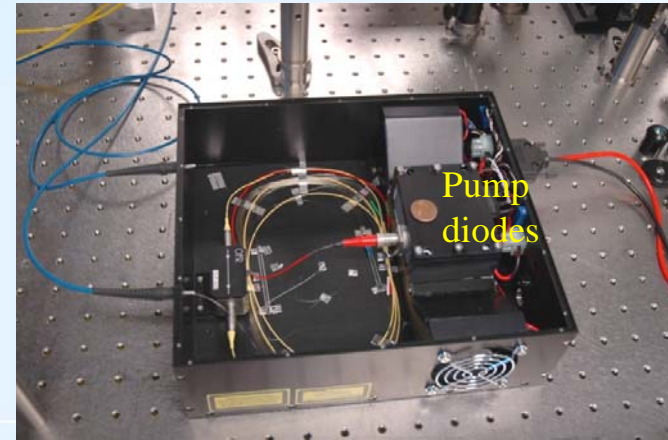
# Laser transmitters (CO2 and O2): Diode Seed Lasers -> Fiber Amplifiers



Tunable diode seed lasers



An example of EDFA technology (Lucent)



## High Reliability 49 dB Gain, 13W PM Fiber Amplifier at 1550 nm with 30 dB PER and Record Efficiency

P. Wysocki, T. Wood, A. Grant, D. Holcomb, K. Chang, M. Santo, L. Braun, G. Johnson  
*Lucent Technologies, Bell Laboratories, 600 Mountain Avenue, Murray Hill, NJ 07974*

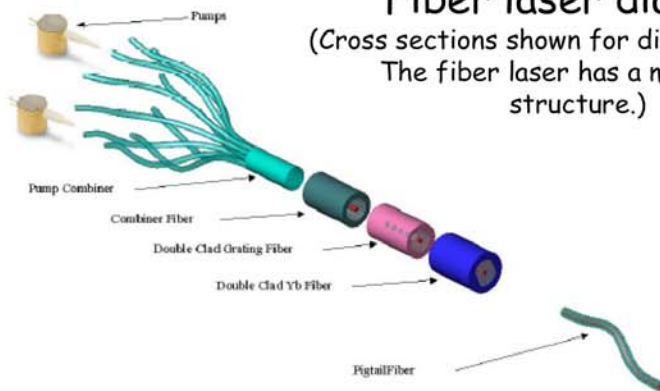
**Abstract:** We present a single-mode, polarization-maintaining, Er/Yb-codoped cladding-pumped amplifier with up to 49 dB of gain, 13W output power, a 30 dB polarization-extinction ratio, 12.9% electrical to optical conversion efficiency, and 42% optical-optical power-stage fiber slope efficiency. This is the most efficient highest PER HPOA ever reported near 1550 nm. The HPOA was built using all fiber-based components qualifiable for long missions in harsh environments.

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OCIS codes: 060.2320, 060.2420, 140.4480

## Characteristics:

- Fiber architecture - permanently aligned
- Closed laser cavity - free from contamination
- Large investments from industry
- Components built to Telcordia standards
- Diode pump technology is very reliable (undersea fiber optic repeaters)
- Distributed thermal load
- Electrical efficiencies: 8- 15%
- Ongoing work for use on satellites
- Wide availability of highly engineered parts
- Wavelength flexibility

## Fiber laser diagram (Cross sections shown for display purposes. The fiber laser has a monolithic structure.)





# Example - Pulsed operation of Large Mode Area EDFA

- Aculight developed an EDFA fiber amplifier using a large mode area (LMA) fiber
  - Single frequency measurements (to date) show:
    - 360 watts pulsed peak power at 1540 nm with no SBS
    - Beam quality:  $M2 < 1.1$
    - Frequency doubling to 770 nm with 56% efficiency
- Calculations show > 500W peak is feasible (space need).

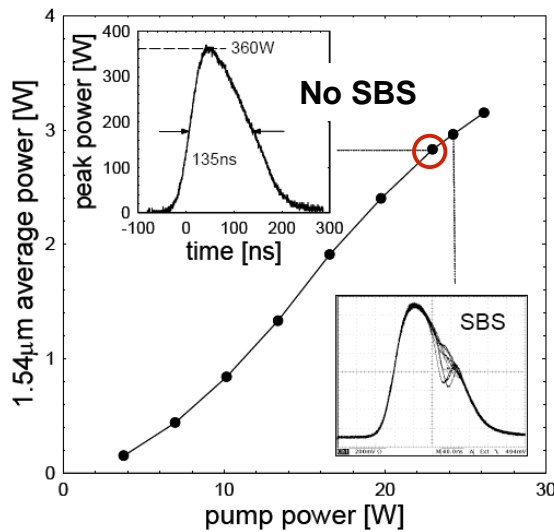


Fig. 3 - 1.54 μm output power at 55 kHz as a function of applied pump power. Additionally, a temporal trace (upper left) with the maximum SBS free peak power and a (long persistence) oscilloscope trace (lower right) showing the onset of SBS are displayed in two separate inserts. The time base of the oscilloscope display was set to 40ns/div.

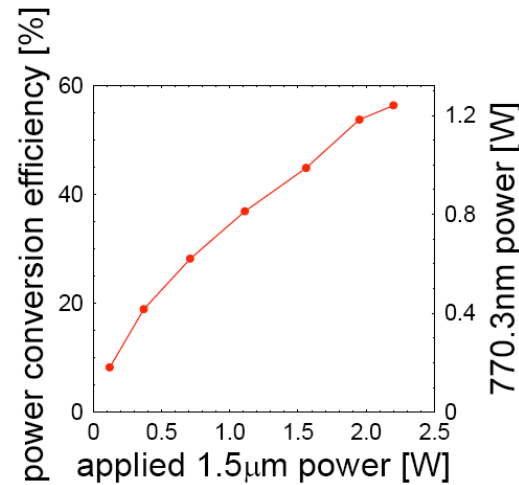


Figure 2-11 Power conversion efficiency to 770.3nm as a function of applied 1540.6nm power.

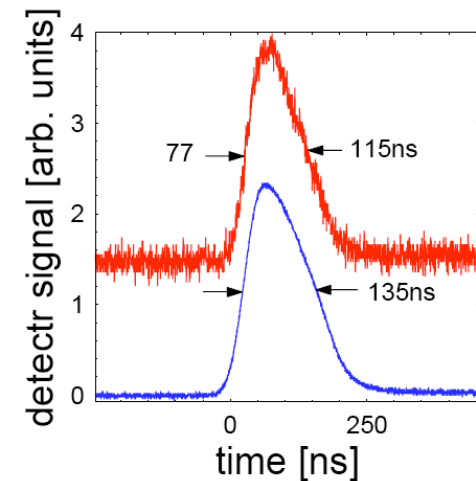
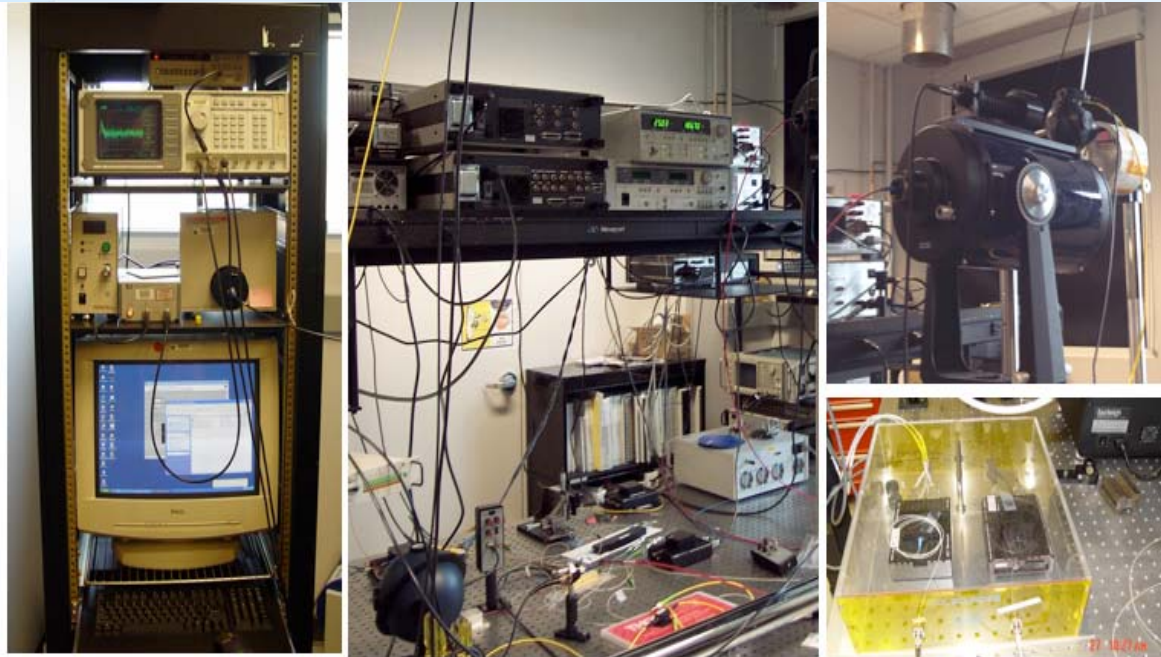


Figure 2-12 Temporal trace of maximum SBS free pulse at 1540.6nm (lower trace) and corresponding frequency doubled pulse at 770.3nm (upper trace).



# CO<sub>2</sub> - Breadboard & Open Path Atmospheric Measurements across Outdoor paths (206 & 400 m one-way)

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Breadboard Co<sub>2</sub> laser sounder Characteristics for March/April 2007

Parameter	Value
CO <sub>2</sub> Line wavelength	1572.33 nm
Measurement Approach	FM-CW (chopped)
Wavelength Scan width	3 Co <sub>2</sub> lines: 650 pm
Receiver Measurements/freq scan	80 GHz
Wavelength scan resolution	1024
Wavelength scan rate	0.6 pm, 80 MHz
Laser beam divergence	0.3 - 1 KHz (varies)
Laser beam power	2.5 mrad
Laser Chopping Duty cycle	300 mW
Laser Chopping Cycle Rate	50%
Test Path length	207 Hz
Telescope Diameter	405 m
Detector type (presently)	20 cm
Detector Bandwidth	InGaAs PIN
Receiver sample rate	300 KHz
Receiver recording duty cycle	400 & 200 KHz
Number of scans/measurement	10%
Dwell Time/measurement	2000
Signal processing approach	65 seconds
	Sample, normalize, equalize, line fit



4/28/08

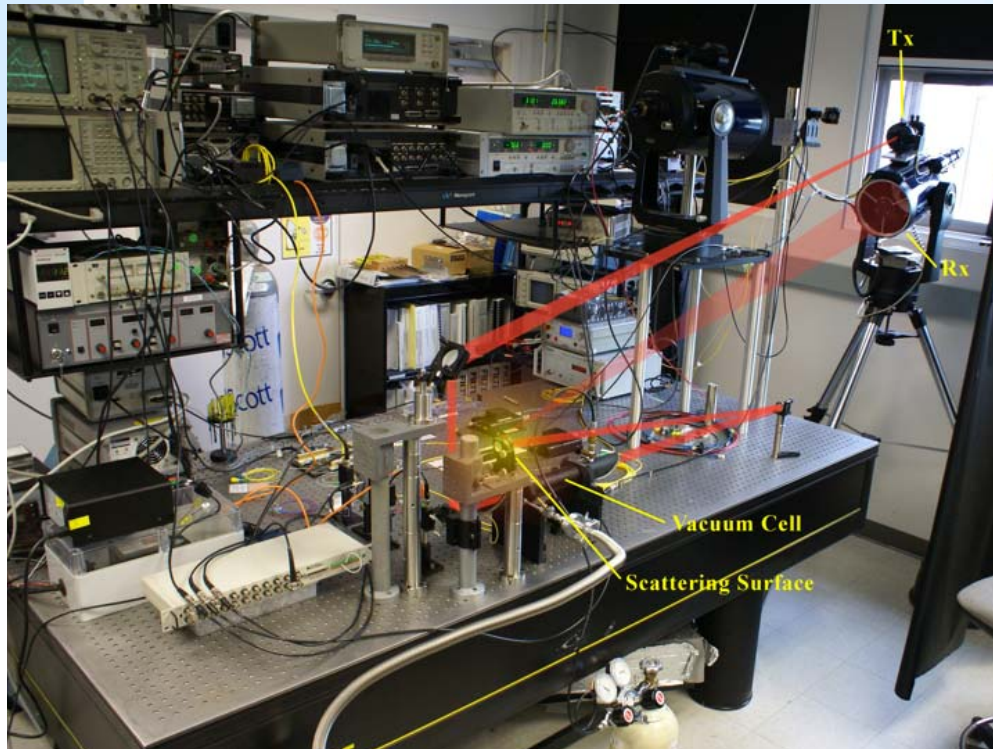
CO<sub>2</sub> Laser Sounder for ASCENDS - NASA Carbon Cycle Workshop

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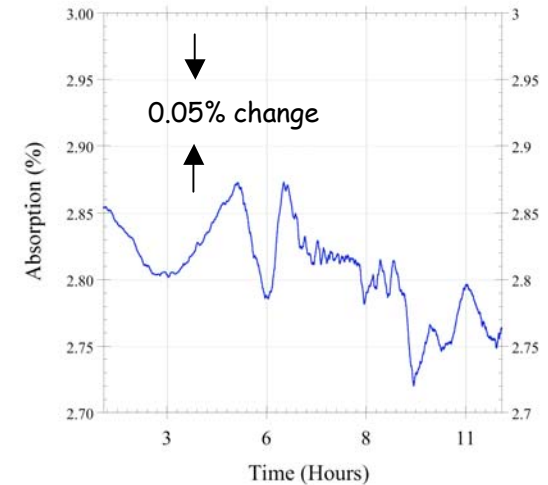
# CO2 Measurement - Stability Demonstration

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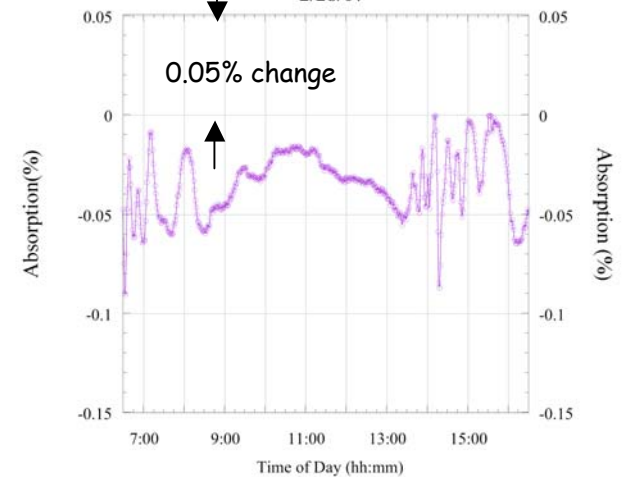


- Must maintain stability between calibrations.
- For space , many hours between calibrations to ground sensors
- Changing etalon fringes can cause drifts in Laser spectrometers
- Goal: keep breadboard drifts to  $< 10^{-3} * \text{expected absorption}$ .
- For 50% absorption => drift errors  $< 0.05\%$

Stability assessment in Lab  
Measurements of Peak Absorption vs Time  
in the lab through a 30 cm long Reference Cell  
with ~10 mbar Co2



Stability of offset (Zero point)  
Measurement in a 30 cm absorption cell  
in the lab with no Co2  
2/28/07

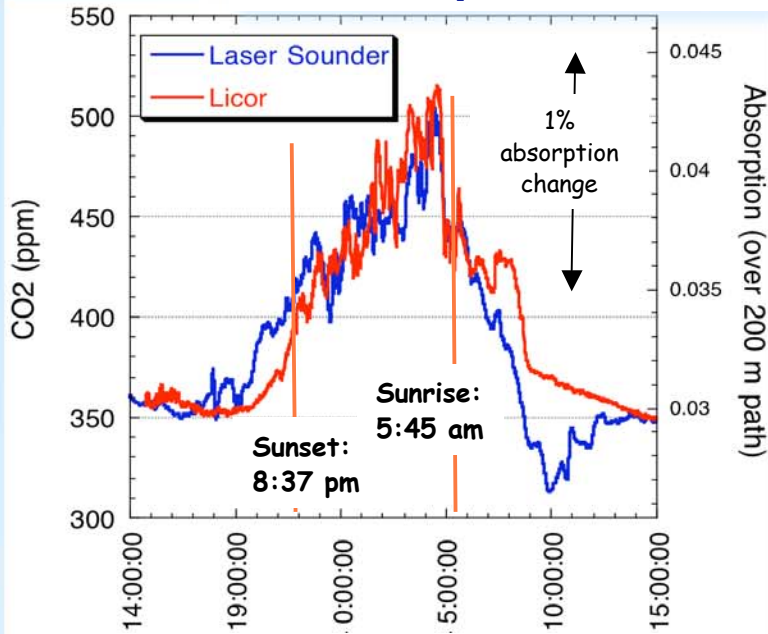


- measurements with 30 cm absorption cell show ~ 0.05% stability for hours

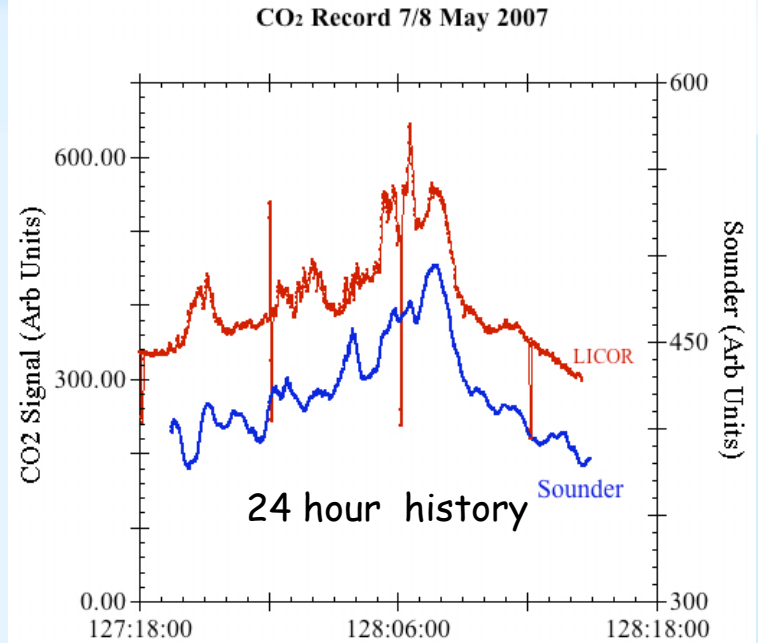


# CO<sub>2</sub> measured over 206 & 405 m open paths Comparison to samples from end point in-situ sensor

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25 hour history



24 hour history

## Earlier Summer Measurements: 206 m path (vegetation active)

- Breadboard measurements offset and scaled
- Show diurnal change in Co2 near surface
- Agreement to 1: 500 in absorption over 1st 16 hrs
- Close to performance needed for space mission
- Improvements later improved reproducibility

### In-situ samples:

Single-point  
measurements (Licor)  
from air intake on B33  
rooftop





# CO<sub>2</sub> Receiver - Photon Counting Detectors

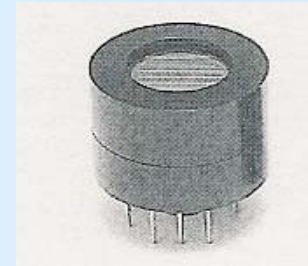
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*Now:*

## Hamamatsu H9170-75 PMT

-Detector used in breadboard receiver

- Turn-key operation
- QE = 9% at 1572 nm
- InP/InGaAs photocathode
- Photocathode: ~5 mm diameter
- Dark count rate ~200 KHz at -80 C (TEC cooled)
- PMT power consumption ~150mW



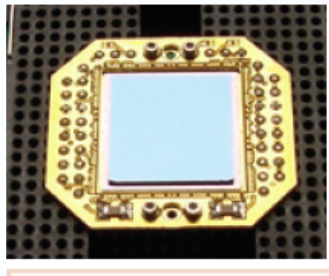
TO-8 PMT package with transmissive photocathode



HV supply and PMT housing

*Under Development:*

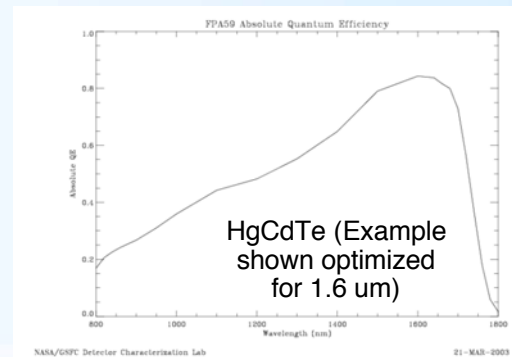
## Adapting HgCdTe Photon Counting Detectors



### The Infrared Detectors for the Wide Field Camera 3 on HST

#### ABSTRACT

We present the performance of the IR detectors developed for the WFC3 project. These are HgCdTe 1Kx1K devices with cutoff wavelength at 1.7 $\mu$ m and 150K operating temperature. The two selected flight parts, FPA#64 (prime) and FPA#59 (spare) show quantum efficiency higher than 80% at  $\lambda=1.6\mu$ m and greater than 40% at  $\lambda>1.1\mu$ m, readout noise of ~25 e<sup>-</sup> rms with double correlated sampling, and mean dark current of ~0.04 e<sup>-</sup>/s/pix at 150K. We also report the results obtained at NASA GSFC/DCL on these and other similar devices in what concerns the QE long-term stability, intra-pixel response, and dark current variation following illumination or reset.



80% photon detection/accumulation efficiency at 1550 nm with ~0.04 dark counts/sec at 150K

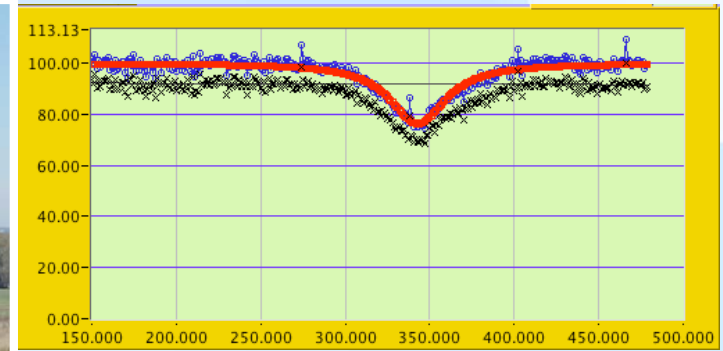
Present read-out configured for photon accumulation



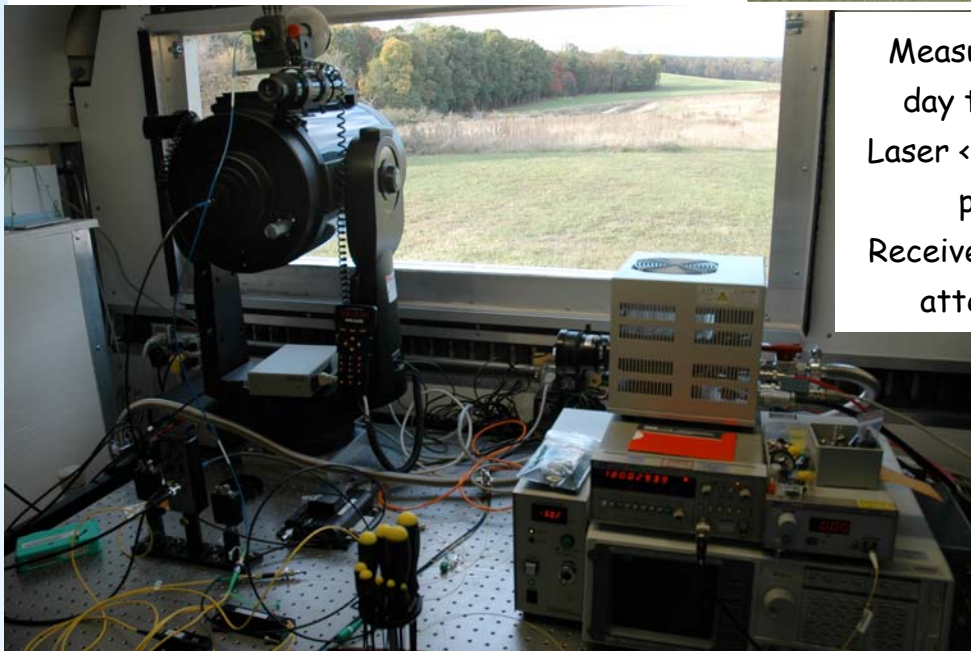


# CO<sub>2</sub> Absorption Measurements from van over 2.2 km path near GSFC using PMT detector

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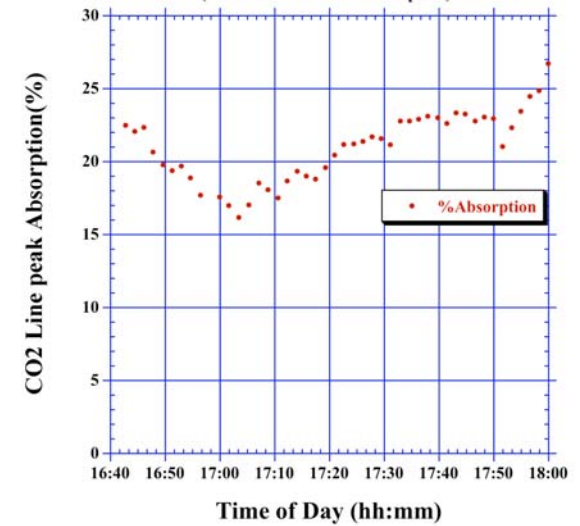


CO<sub>2</sub> Line scan using PMT receiver



Measurements -  
day to sunset  
Laser < 0.4W peak  
power  
Receiver PMT was  
attenuated

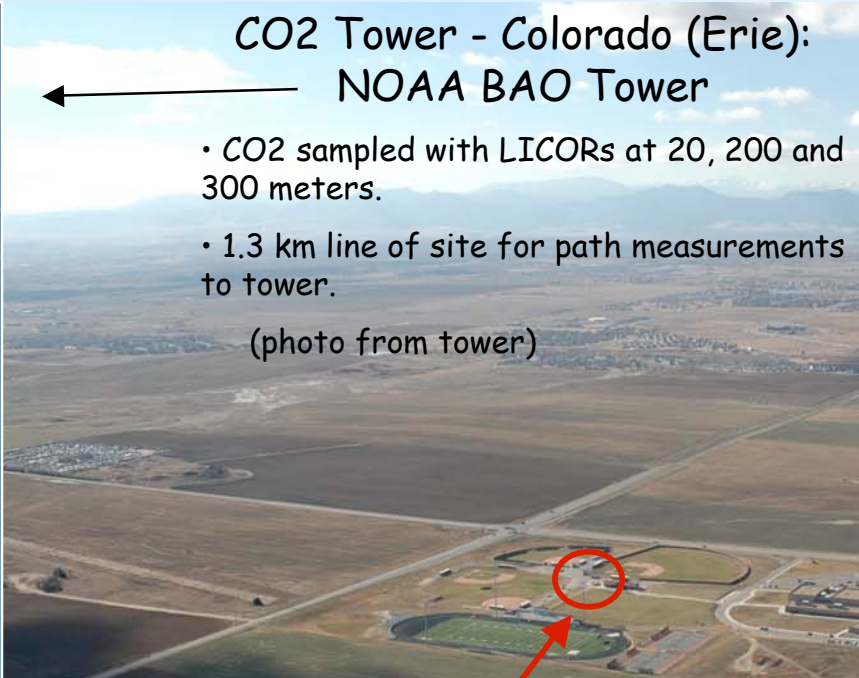
CO<sub>2</sub> Absorption History measured with  
Co<sub>2</sub> Sounder field instrument  
over 2.2 km Optical Path at GSFC  
to side of water tower using PMT detector  
11-3-07  
(100 sec and 10<sup>4</sup> line scans/point)





# CO2 Measurements at NOAA tower, Erie CO (G. Allan, E. Wilson, H. Riris)

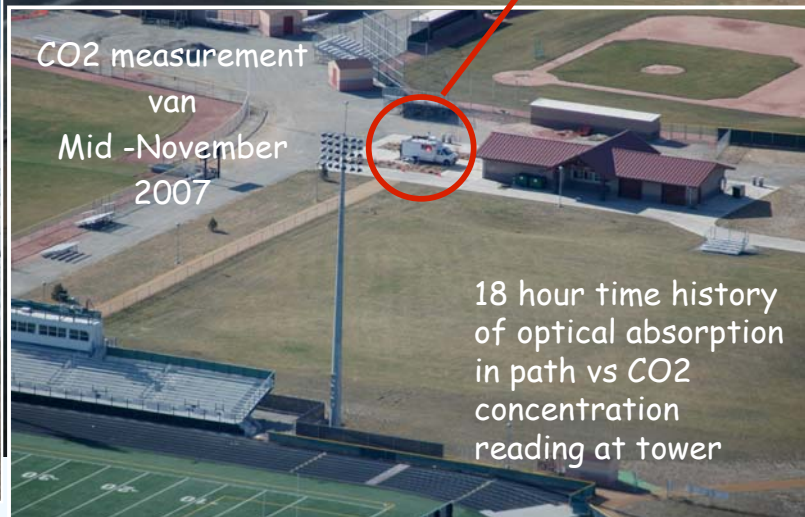
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## CO2 Tower - Colorado (Erie): NOAA BAO Tower

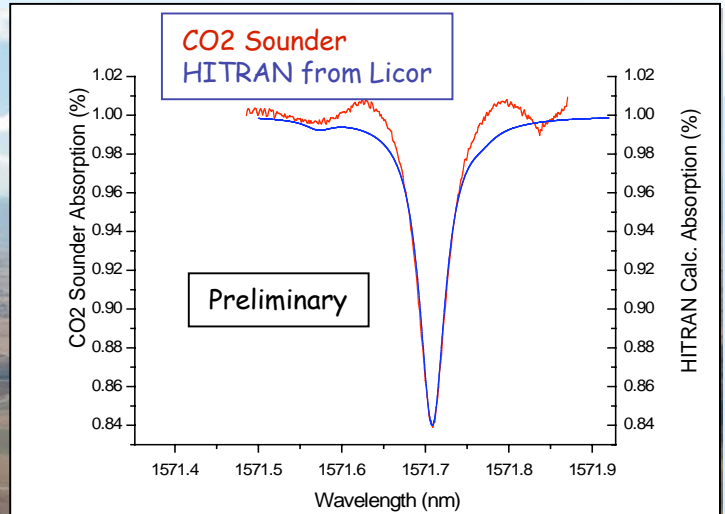
- CO2 sampled with LICORs at 20, 200 and 300 meters.
- 1.3 km line of site for path measurements to tower.

(photo from tower)

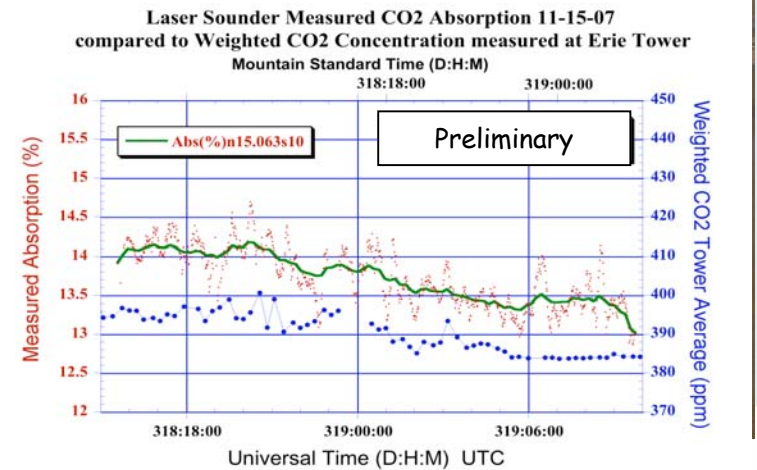


CO2 measurement  
van  
Mid -November  
2007

18 hour time history  
of optical absorption  
in path vs CO2  
concentration  
reading at tower



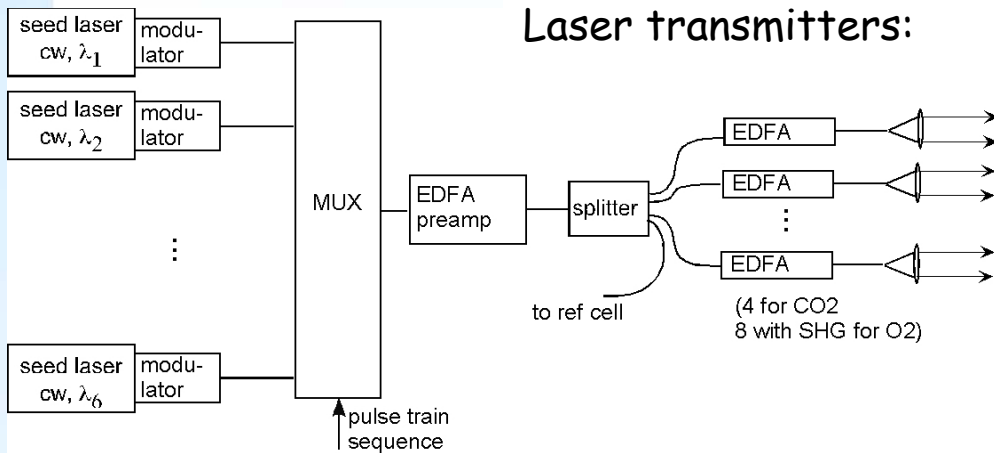
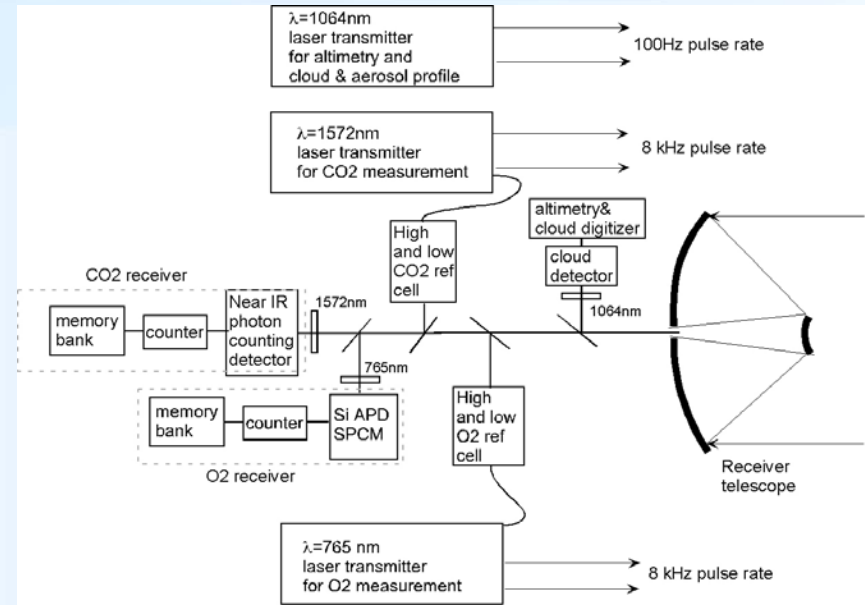
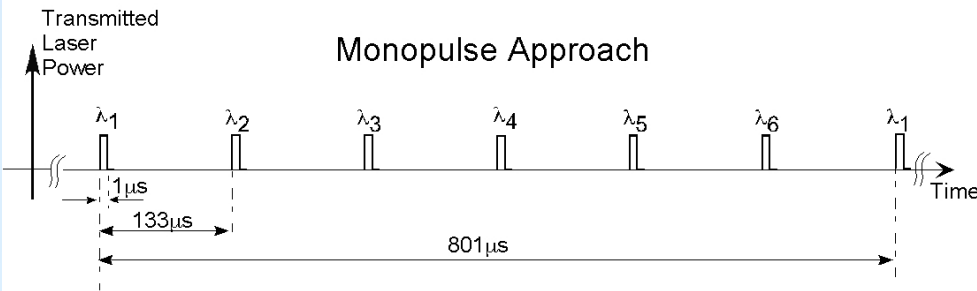
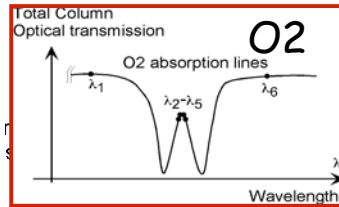
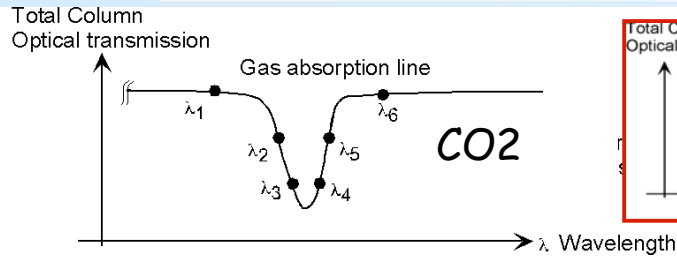
Comparison of measured CO2 line scan with HITRAN Prediction based on the Tower LICOR





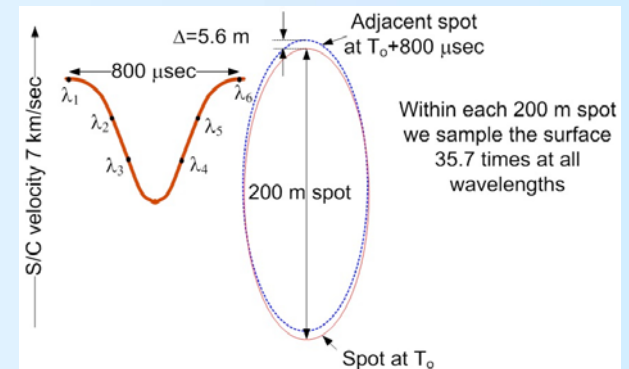
# Measurement approach for Aircraft & Space

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(Each laser spot contains illumination from 35 cycles through the 6 laser wavelengths)

Co-aligned and Overlapped spots on ground

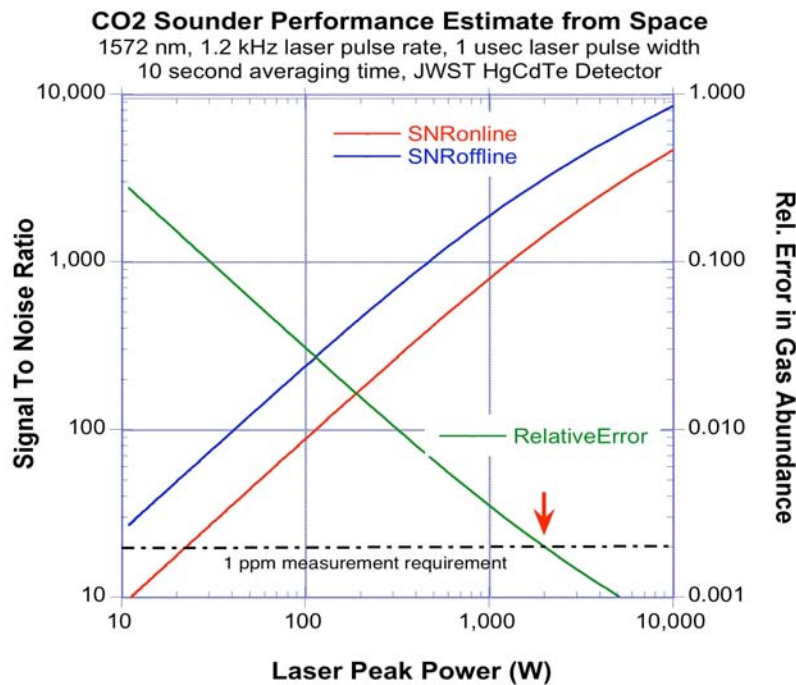




# Space: SNR & Relative Measurement Errors

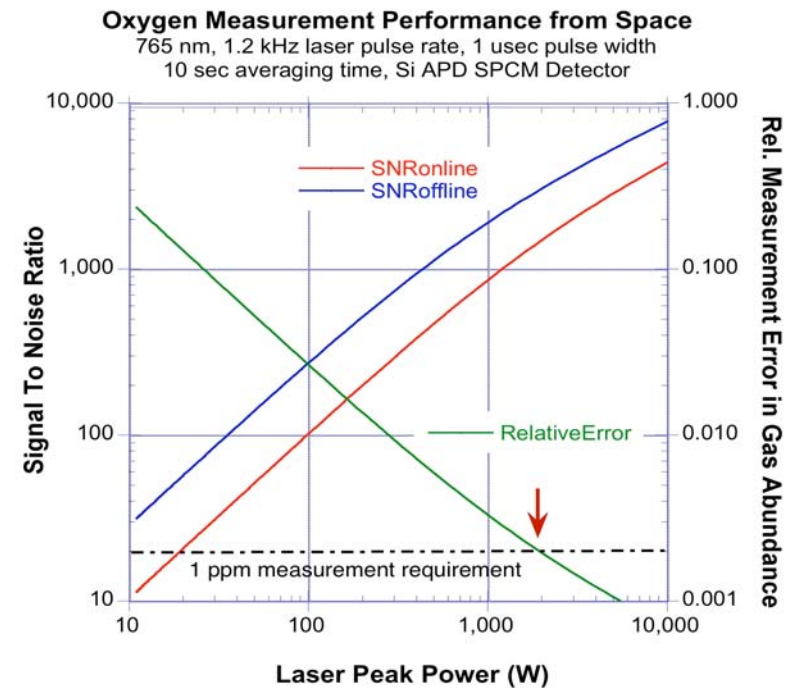
(10 seconds observing time, 500 km orbit, 1.5m telescope)

## CO2 column measurement



4 EDFA's, 500W pk power

## O2 column measurement



8 EDFAs, 500W pk power, 50% doubling

Rel Measurement Errors scale as (peak laser power)<sup>-1</sup> \* (T)<sup>-1/2</sup>



# Next steps - Airborne ASCENDS Precursor for NASA ESTO IIP Program (start in FY09)

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## CO<sub>2</sub> Laser Sounder for ASCENDS Mission – Technology development and airborne demonstrations

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December 12, 2007

Submitted in response to NNH07ZDA001N-IIP

### Summary:

We propose to advance measurement technology and reduce the risk and cost for the ASCENDS mission. The measurements from our targeted laser instrument for space will measure CO<sub>2</sub> column abundance and fluxes with a spatial resolution of ~100 km, and will meet or exceed the science needs as summarized in the mission description.

Our pulsed laser approach measures the energies of laser pulses reflected from the Earth's surface. Laser transmitters for CO<sub>2</sub> and O<sub>2</sub> are rapidly tuned on and off selected atmospheric CO<sub>2</sub> and O<sub>2</sub> absorption lines near 1572 nm and 765 nm. A laser at 1064 nm is used to measure surface height and aerosol backscatter profile. Time gating is used to isolate the echo pulses from the surface and to minimize errors from atmospheric scattering and solar background.

Our proposal leverages strong understanding and capabilities developed over the past 7 years with support from the ESTO ACT and IIP programs. This work addressed many aspects of the measurement, improved the technique and technologies, and successfully addressed a number of significant error sources. We have demonstrated both CO<sub>2</sub> and O<sub>2</sub> measurements in open path tests, and made precise measurements of CO<sub>2</sub> absorptions over many days. Our understanding of CO<sub>2</sub> fluxes is enhanced by co-investigators who conduct research on atmospheric CO<sub>2</sub>.

We propose to advance the readiness of the instrument technologies. We will demonstrate these in an airborne precursor instrument for the ASCENDS mission, to use as a simulator for a space mission, as a test bed for improving components, and for airborne science campaigns. We will also optimize the calibration approaches and flux recovery and improve the fidelity of the space instrument definitions. Our proposed work will advance the technology TRL > 5, to allow an instrument baseline that can be developed for space mission in a ~3-year period.

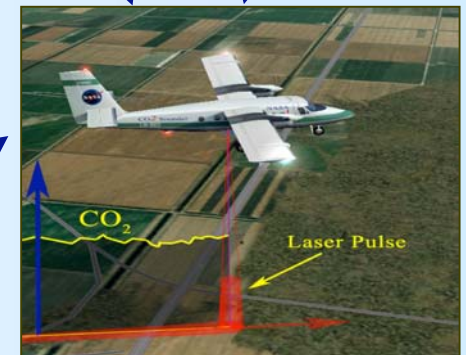
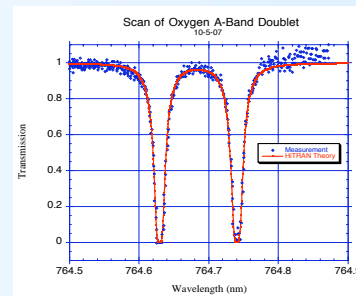
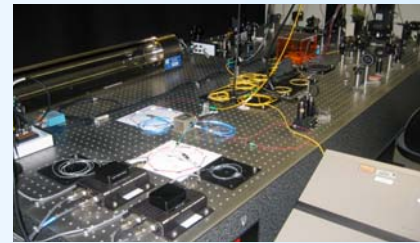
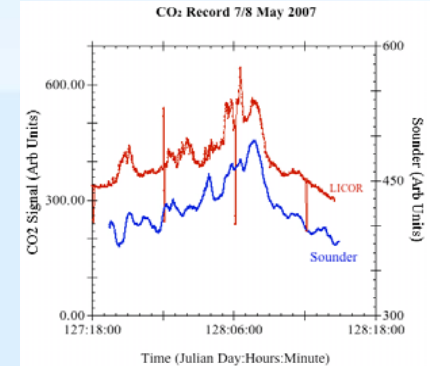
## Airborne Precursor for ASCENDS Mission: Calibrated mixing ratio measurements

Why ?

- Simulator for Space
- Test bed for evolving components
- Airborne science campaigns

CO<sub>2</sub> Laser Sounder for ASCENDS Mission – Technology development and airborne demonstration

ii



ASCENDS Airborne  
Precursor Instrument



# Summary

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## A needed & challenging space

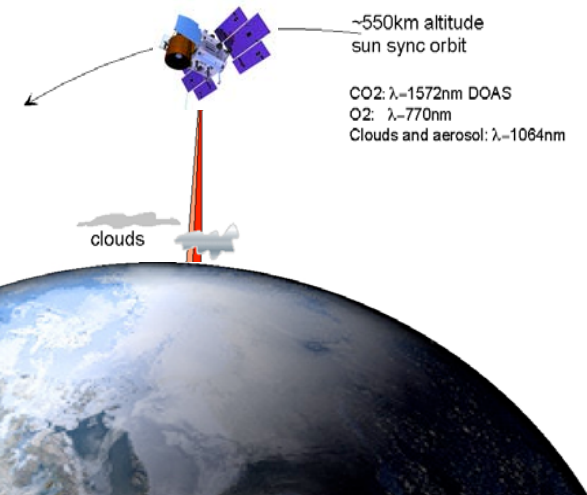
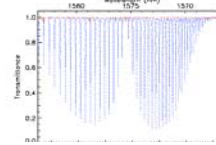
### measurement:

- Complex measurement environment
- Many potential error sources
- Need high precision & accuracy

### Our measurement approach:

- CO<sub>2</sub> at 1572 nm, O<sub>2</sub> at 765 nm
- CO<sub>2</sub> lower tropospheric column
- O<sub>2</sub> total column (surface pressure)
- Altimetry & atmos. backscatter profile
- Temperature & WV from atmospheric models
- Pulsed EDFA lasers
- Direct detection in all channels
- Time gated Photon-counting receiver

Measures:  
- CO<sub>2</sub> tropospheric column  
- O<sub>2</sub> tropospheric column  
- Cloud backscattering profile



### Evolution:

- ✓ Initial concept (2001) & lab measurements
- ✓ Breadboard instrument development
- ✓ Space instrument component need analysis
- ✓ White paper submitted to NRC RFI
- ✓ CO<sub>2</sub> & O<sub>2</sub> spectroscopy & error analysis
- ✓ Field measurements at CO<sub>2</sub> tower (2007)
- CO<sub>2</sub> Airborne demonstrations (fall 2008)
- CO<sub>2</sub> & O<sub>2</sub> airborne precursor (2009-11)